

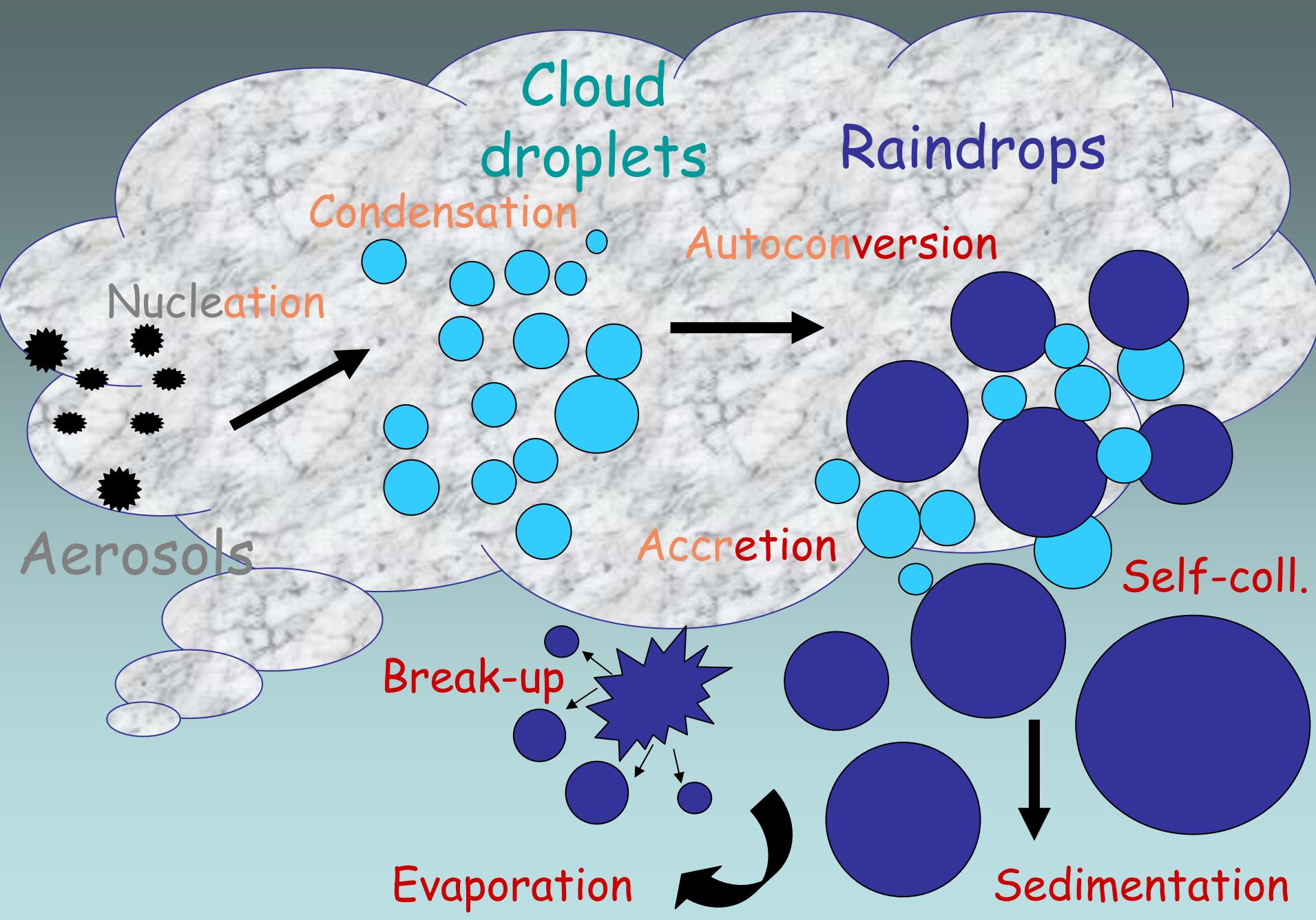
Warm microphysical schemes available in MesoNH

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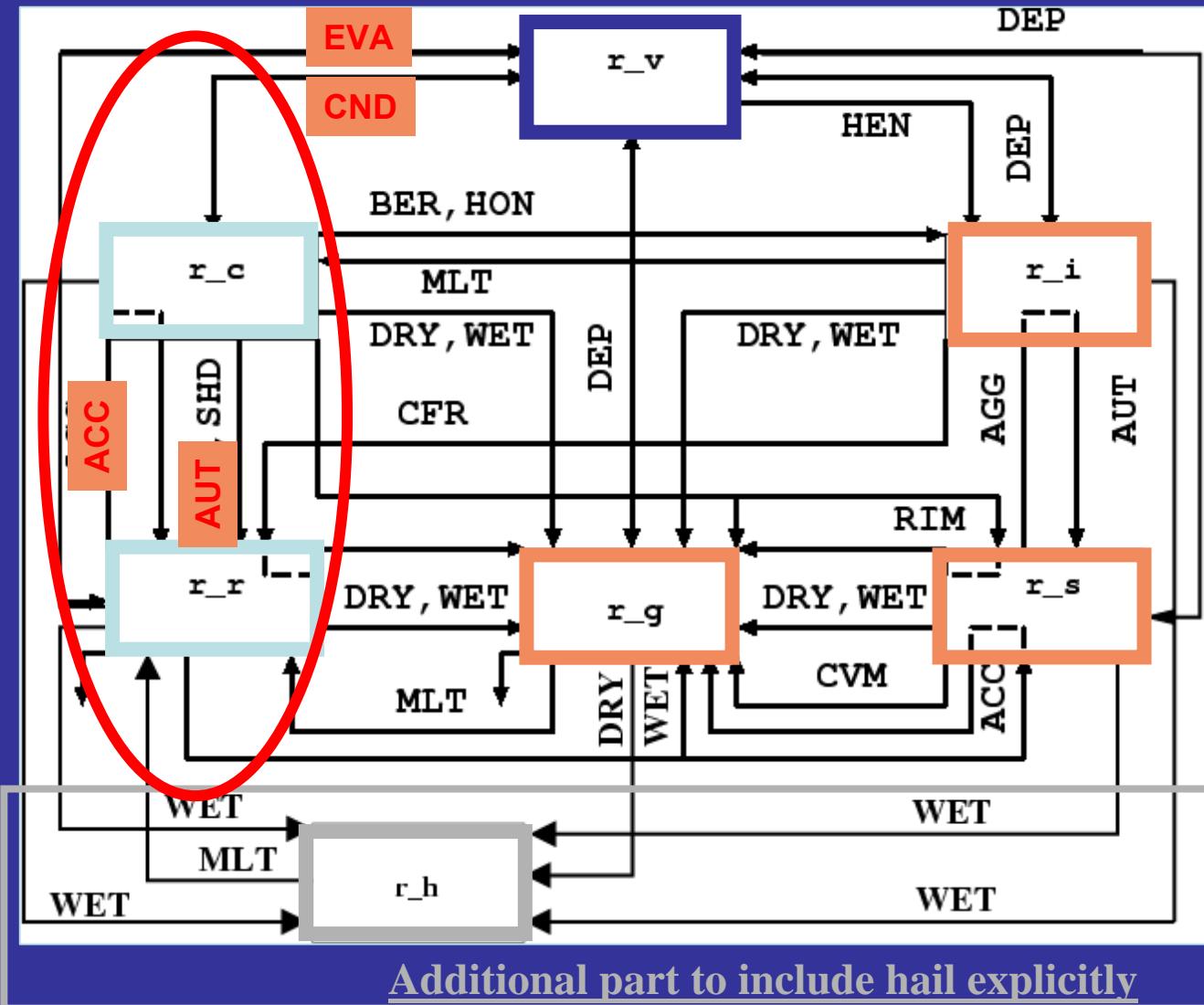
« Warm » microphysics ?

- Key-processes to follow the evolution of the cloud and rain **mixing ratios** (kg of water/kg of dry air)
- **2-moment schemes** can also predict the cloud and the rain number concentrations (link with CCN)
- **Bulk** versus **Spectral (bin)** microphysics ($50 < \text{bins} < 100$)
- Current **assumptions** in the schemes:
 - spherical drop-droplets of size D ($m = aD^3$),
 - fall speed $V(D) = cD^d (\rho_{00}/\rho)^{0.4}$,
 - exponential, lognormal or gamma size-distributions $n(D)$ (bulk schemes)



Microphysical Scheme diagram

Corresponds to a reduced set of a full mixed-phase microphysical scheme !



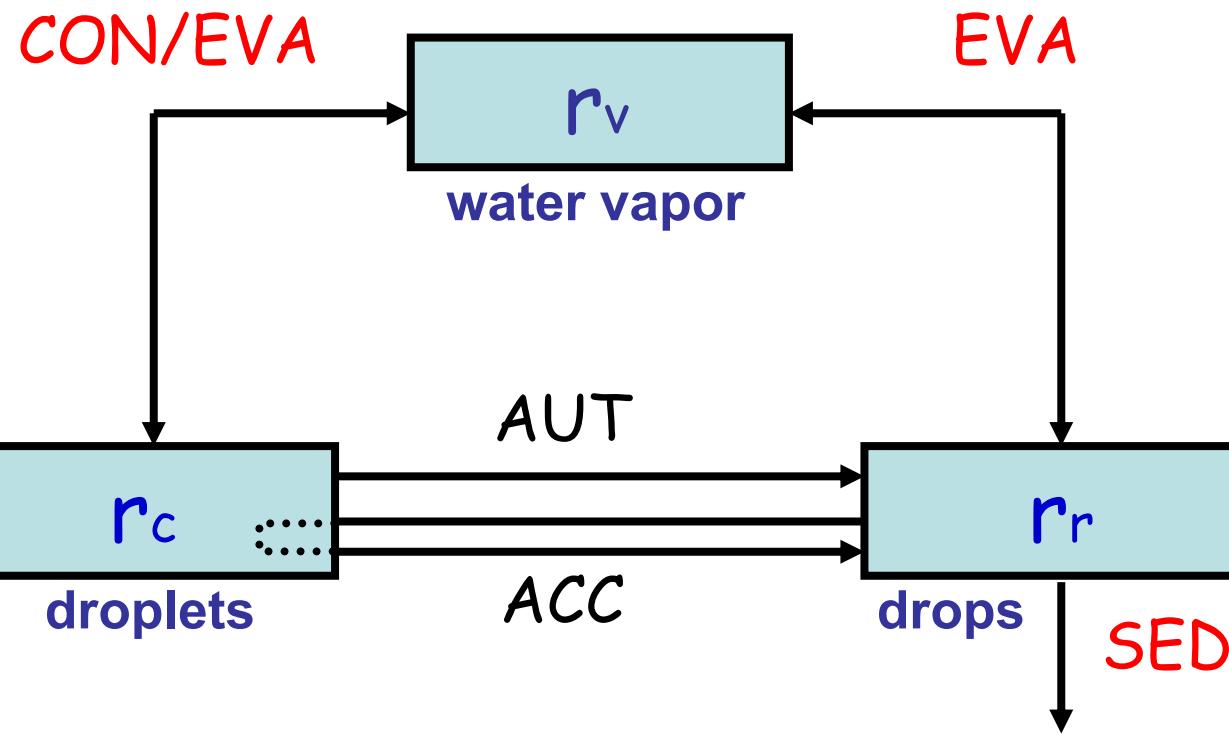
3 bulk schemes are available

- The “**Kessler**” scheme (Kessler, 1969) is the simplest scheme used to simulate warm rain precipitation events {r_cloud, r_rain}
- The “**C2R2**” 2-moment scheme (Cohard & Pinty, 2000) includes the evolution of the droplet/drop concentrations {N_cloud, N_rain}
- The “**KHKO**” scheme is an adaptation of (Khairoutdinov & Kogan, 2000). It is a parametric scheme to simulate the drizzle

No spectral scheme yet !

Some merits of the Kessler scheme

- Simplicity (5 processes)



- **CON/EVA**
adjustment to saturation with conservation of $h = C_p T + Lq_{vs}(T)$
- **EVA** explicit drop evaporation rate
- **SED** explicit differentiation of sedimentation fluxes

Autoconversion in Kessler scheme

- Autoconversion: cloud droplets → rain drops

$$(\partial \rho_a r_r / \partial t)_{AUT} = K \ Max(0, \rho_a r_c - \rho_a \tilde{r}_c)$$

with

$$K = 10^{-3} s^{-1}$$

$$\rho_a \tilde{r}_c = 0.5 g/m^3$$

- Many variants: Manton and Cotton (1977) or Liu and Daum (2006) could be used to relate the two adjusted parameters to the droplet spectra

Accretion in Kessler scheme

- Accretion: raindrop+cloud droplets → bigger raindrop

$$(\partial \rho_a r_r / \partial t)_{ACC} = \int_0^\infty r_c \left(\frac{\pi}{4} D^2 V(D) \right) E n_r(D) dD$$

with

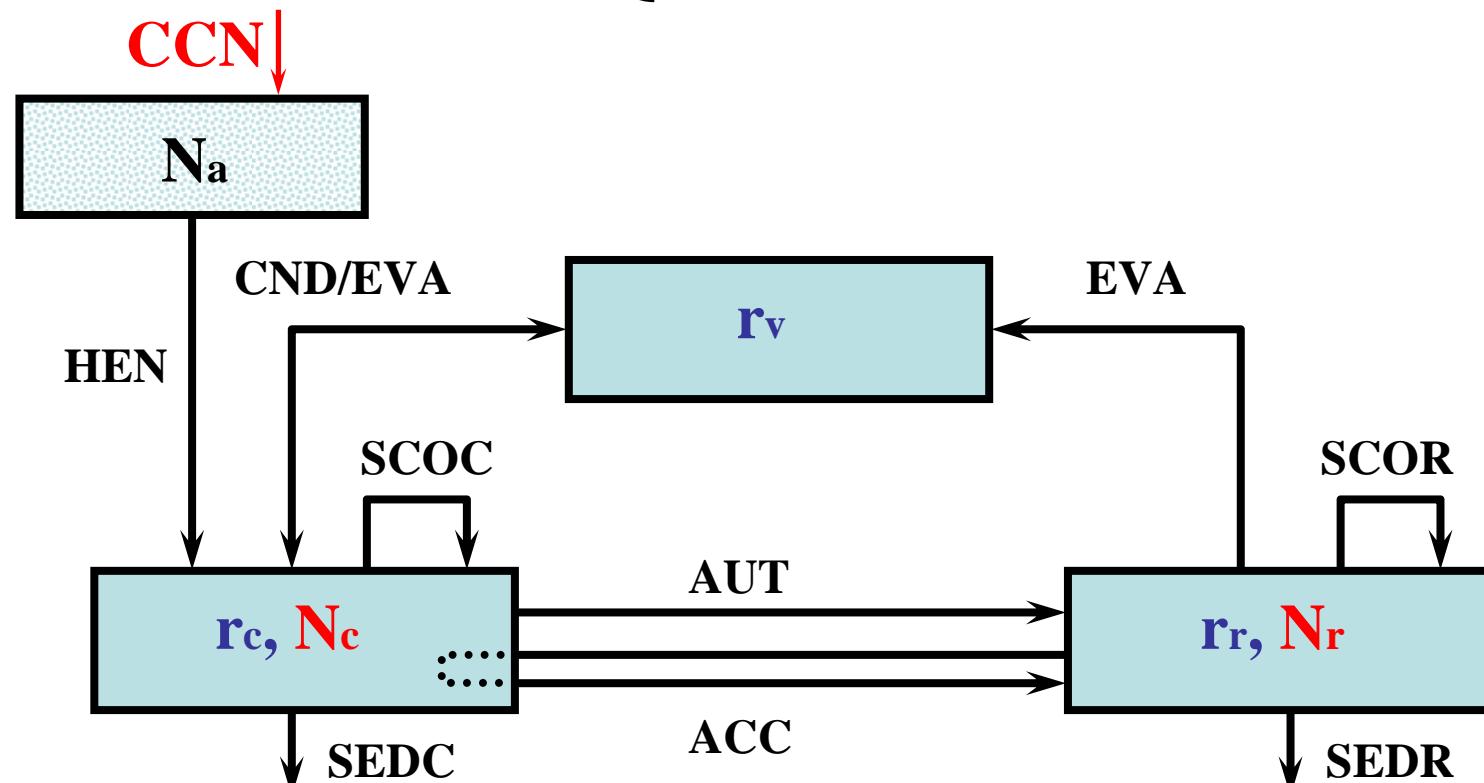
$$E = 1$$

- Exact integration for a given $n_r(D)$ size distribution but uncertainties about the value of the mean collection efficiency E

The 2-moment C2R2 scheme

Generalized gamma size distribution law

$$\left\{ \begin{array}{l} n(D) = N \frac{\alpha}{\Gamma(\nu)} \lambda^{\alpha\nu} D^{\alpha\nu-1} \exp(-(\lambda D)^\alpha) \\ \alpha, \nu \text{ are free parameters} \end{array} \right.$$



Nucleation in C2R2 scheme

Theory of Twomey

- Activation spectrum :

$$N_{CCN}(s_{v,w}) = \int_0^{s_{v,w}} n(s) ds$$

- Sursaturation :

$$\frac{ds_{v,w}}{dt} = \psi_1(T, P)_w - \psi_2(T, P) \frac{dr_c}{dt}$$

- Diffusion growth (Köhler) :

$$D \frac{dD}{dt} = 4G(D, T, P)(s_{v,w} - y(D, T))$$

After combination → $s_{v,wMAX}$

$$\psi_1(T, P)_w \approx \psi_2(T, P) \frac{dr_c}{dt} \quad \text{and}$$

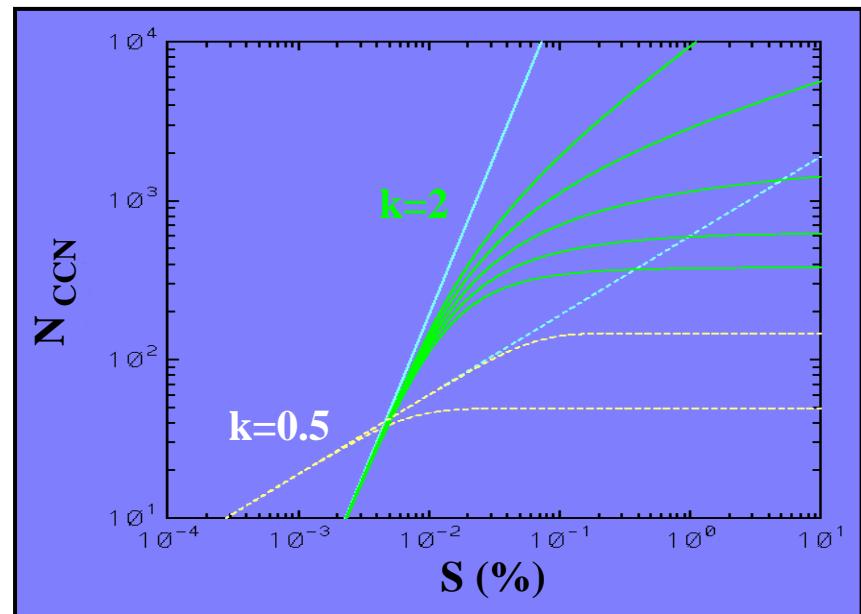
$$\frac{dr_c}{dt} \approx \psi_3(D, T, P) \int_0^{s_{v,wMAX}} n(s') (s_{v,wMAX}^2 - s'^2)^{1/2} ds'$$

→ Number of activated CCN

Twomey (1959) : $N_{CCN} = Cs_{v,wMAX}^k$

Cohard *et al.* (1998) :

$$N_{CCN} = Cs_{v,wMAX}^k F\left(\mu, \frac{k}{2}, \frac{k}{2} + 1, -\beta s_{v,wMAX}^2\right)$$



... calibration of N_{CCN} from aerosols

Autoconversion in C2R2 scheme

- **Initiation regime:** w/o raindrops

Berry and Reinhardt (1974) parameterization
recently reviewed by Gilmore and al. (2007)

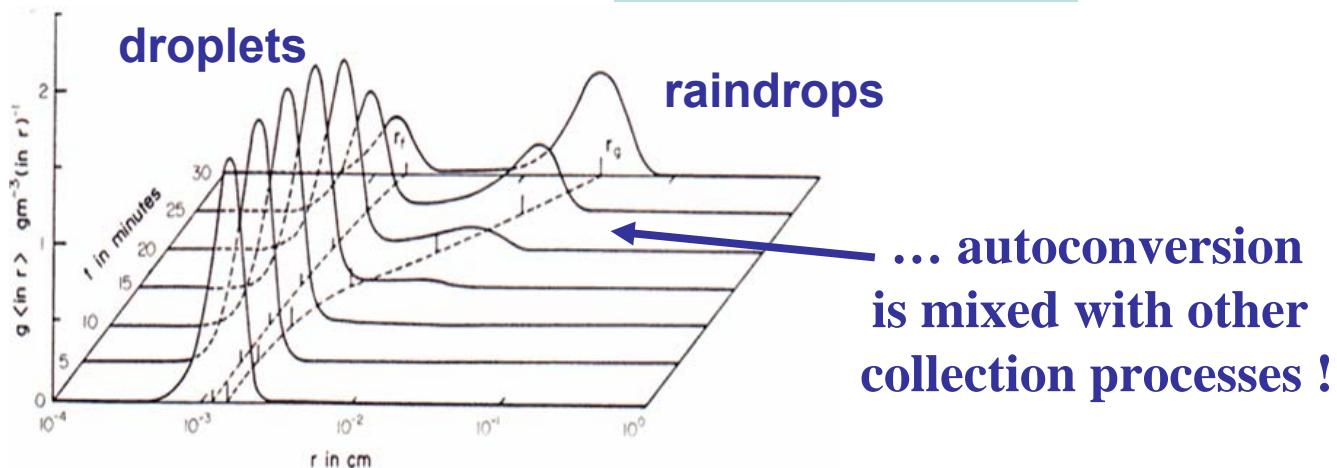
$$\left\{ \begin{array}{l} \left(\frac{\partial r_r}{\partial t} \right)_{AUT} = - \left(\frac{\partial r_c}{\partial t} \right)_{AUT} = \frac{L}{\rho_a \tau} \\ \left(\frac{\partial N_r}{\partial t} \right)_{AUT} = -3.5 \times 10^9 \frac{\rho_a L}{\tau} \end{array} \right.$$

$$\tau : \text{time needed to initiate the rain spectrum} \quad \tau = \frac{1}{\rho_a r_c} (0.5 \times 10^6 \sigma_c - 7.5)^{-1}$$

$$L : \text{rainwater content at } \tau \quad L = 2.7 \times 10^{-2} \rho_a r_c (1/16 \times 10^{20} \sigma_c^3 D_c - 0.4)$$

- **Feeding regime:** with raindrops

$$\left(\frac{\partial N_r}{\partial t} \right)_{AUT} = \frac{N_r}{r_r} \left(\frac{\partial r_r}{\partial t} \right)_{AUT}$$



Collection in C2R2 scheme

Analytical solution of the Stochastic Coalescence Equation (SCE)

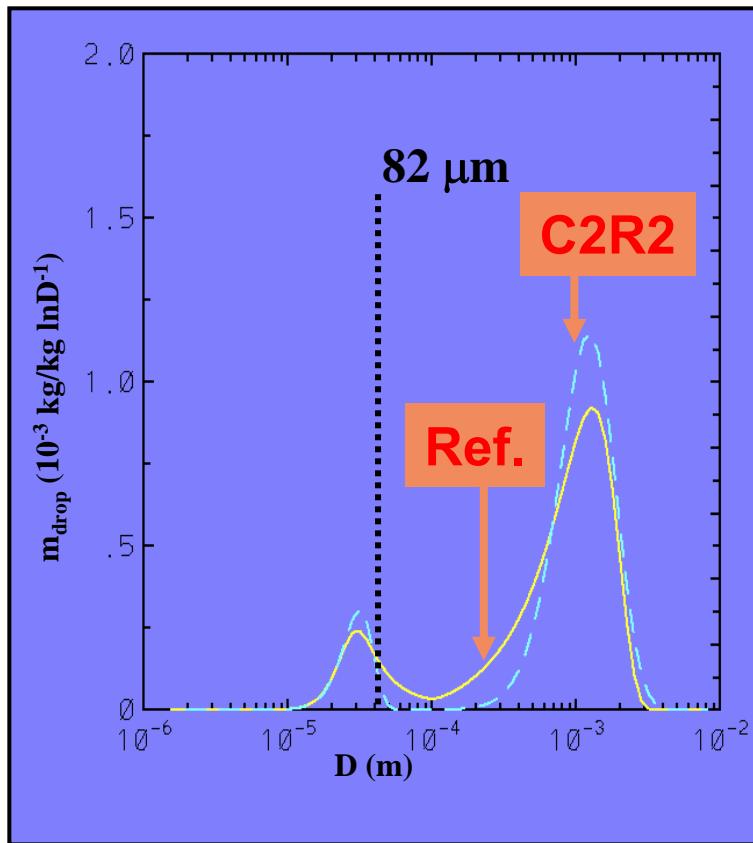
$$\frac{dn(x,t)}{dt} = 0.5 \int_0^x K(y, x-y) n(y, t) n(x-y, t) dy - n(x, t) \int_0^\infty K(y, x) n(y, t) dy$$

with $n(x) = n_c(x) + n_r(x)$ and $K(x, y)$: polynomial kernels of Long

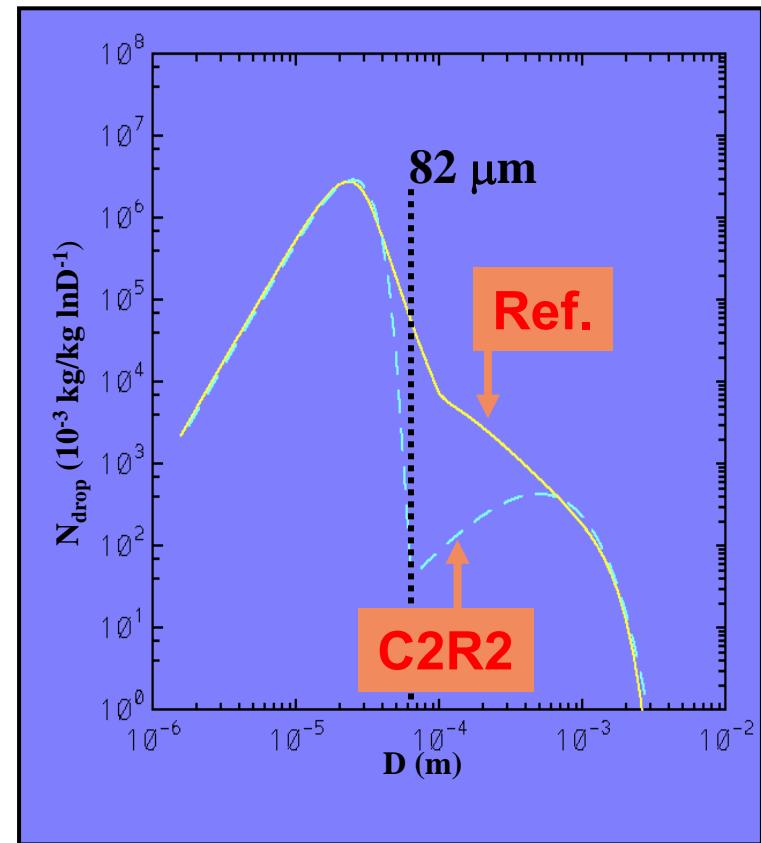
$$\begin{aligned} \frac{dn(x,t)}{dt} &= SCE(n_c) + SCE(n_r) + \int_0^x K(y, x-y) n_c(y, t) n_r(x-y, t) dy \\ &\quad - n_r(x, t) \int_0^\infty K(y, x) n_c(y, t) dy - n_c(x, t) \int_0^\infty K(y, x) n_r(y, t) dy \\ &= \underbrace{SCOC(n_c) + SCOR(n_r)}_{\text{Self-Collection}} + \underbrace{ACC(n_r) + ACC(n_c)}_{\text{Accretion}} \end{aligned}$$

C2R2 versus spectral scheme

Mass size-distributions



Number size-distributions



... an acceptable agreement !

The parametric KJKO-like scheme

Similar to C2R2 scheme but adapted to the drizzle

Autoconversion

$$\left(\frac{\partial r_r}{\partial t}\right)_{AUT} = 1350 r_c^{2.47} N_c^{-1.79}$$

$$\left(\frac{\partial N_r}{\partial t}\right)_{AUT} = \left(\frac{\partial r_r}{\partial t}\right)_{AUT} \frac{4\pi\rho_w}{3\rho_a} r_0^3$$

with $r_0 = 25 \mu m$

Ajustment from
LES and spectral
microphysics

No threshold of
the mixing ratio
=/= Kessler and
C2R2 schemes !!!

The parametric KJKO-like scheme

Similar to C2R2 scheme but adapted to the drizzle

Accretion

$$\left(\frac{\partial r_r}{\partial t}\right)_{ACC} = 67(r_c r_r)^{1,15}$$

Ajustment from
LES and spectral
microphysics

$$\left(\frac{\partial N_r}{\partial t}\right)_{ACC} = \frac{N_c}{r_c} \left(\left(\frac{\partial r_r}{\partial t}\right)_{ACC} + \left(\frac{\partial r_r}{\partial t}\right)_{AUT} \right)$$

Sedimentation

Fallspeeds of r_r and N_r are adjusted

In conclusion ...

- Bulk microphysical schemes can be used to simulate rainfall and drizzle but C2R2 and KHKO schemes are more suitable for BL clouds
- Autoconversion is a key-process to initiate a rain spectrum from cloud droplets (particularly important for BL clouds)
- Microphysics needs a detailed description of BL cloud spectra (mixing ratio AND number concentration) → importance of CCN nucleation treatment to simulate the CCN spectra variability
- Strong impact on radiative transfer (SW and LW)