



A 3D high resolution model for local fog prediction

Rachel Capon

Yongming Tang, Pete Clark, Richard Forbes

Met Office Joint Centre for Mesoscale Meteorology

Introduction – Radiation Fog

Why and Why is it so Difficult to Forecast?
Local Stand-Alone Model Concept

Unified Model at Very High Resolution

Model Physics

Radiation Temperature Profile
Stable Boundary Layer

Single Profile Forcing

Conclusions

- Fog and low visibility conditions can have a high impact on both aviation and ground transport.



CNN

Luxair crash 06.11.2002 – 18 dead



BBC

Accident on M4
near Cardiff
10.12.2003

After the evening transition, the energy balance in the lower atmosphere is no longer dominated by incoming solar radiation.

Following evening transition:

- Turbulent diffusion of moisture and possible dew deposition

- Radiative cooling
 - of surface
 - AND
 - of air (mostly to surface)

“The development of radiation fog is primarily controlled by a balance between radiative cooling, which encourages fog, and turbulence, which inhibits it.”

Roach et al 1976, Brown and Roach 1976, Turton and Brown 1987

- Radiative cooling (of surface and air) in lower atmosphere.
- Increasing stability near to surface.
- Damping of turbulent diffusion, dew deposition inhibited.
- If the air is sufficiently moist, cooling leads to saturation and condensation of water droplets forming fog!
- Wind lulls coincide with maximum cooling and significant fog development.

Radiation Fog – DEVELOPMENT and DISSIPATION



DEVELOPMENT

- Fog droplets also cool radiatively and grow.
- Once the fog is developed, the surface inversion moves to the top of the fog layer.
- Radiation, cooling and condensation now take place mainly at the fog top.

DISSIPATION

- Fog droplets settle due to gravity.
- Cold thermals occur in the fog and entrain warm air. The fog becomes unstable, behaving like Strato-Cumulus.
- At sunrise the fog may burn off if optically thin or near its edges.

Numerical fog forecasting is difficult!

- Fog evolution depends sensitively on several physical processes.
- These processes occur on a sub-grid scale. Hence they must be parametrised.

Radiation

Stable Boundary Layer – turbulence

Surface Exchange

Soil

Microphysics

Numerical fog forecasting is difficult!

- Orography is also important – it can cause pooling of cold air. A 3D model is needed to represent this and the winds.
- LW radiative cooling is sensitive to [synoptic] cloud cover.
- Vertical resolution is important, e.g. to capture radiative fluxes near the surface. Fog may be very shallow, only 2-5m deep, in the initial stages!

- Centrally run NWP at ~1 km scale over areas the size of UK will be feasible by end of the decade.
- Communication of large volumes of output data to the field is likely to remain a major problem.
- Develop a model (based on UM) to run at high resolution over a small domain, forced by a single profile (model or sonde) – to be run on a local PC in the field.

Met Office
Supercomputer



GLOBAL/
MES

— — — — — →
Communications

Local 'PC'



Unified Model at Very High Resolution

1 km horizontal grid, 76 vertical levels

Model Physics

Optimise for fog forecasting at very high resolution

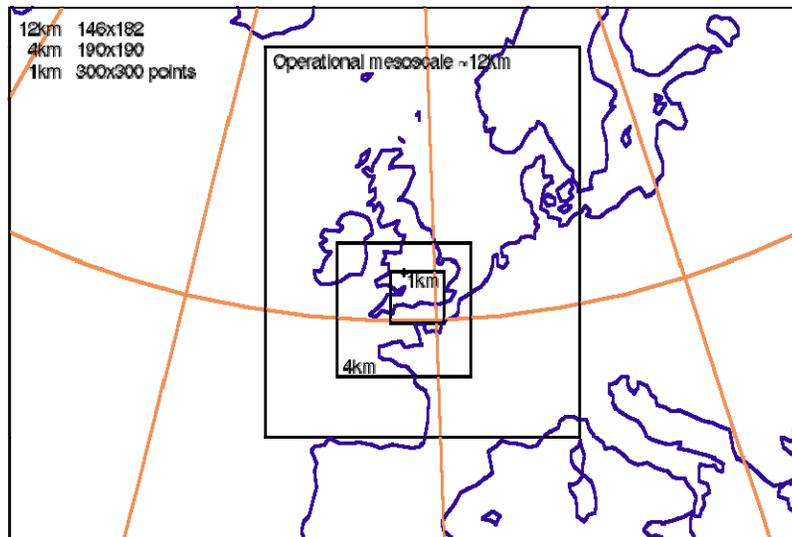
Single Profile Forcing



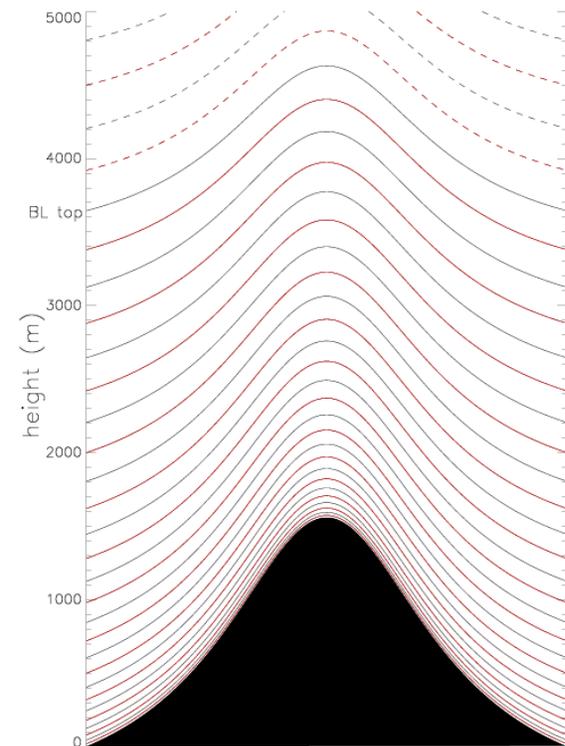
Very High Resolution Modelling

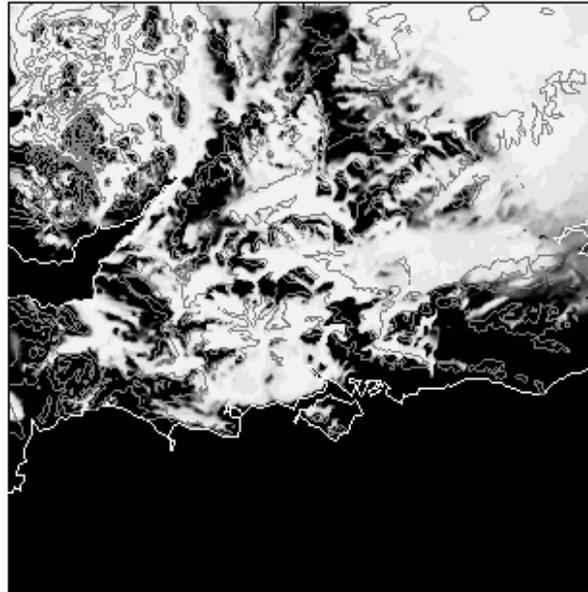
1-Way Nested Limited Area Configuration

HORIZONTAL



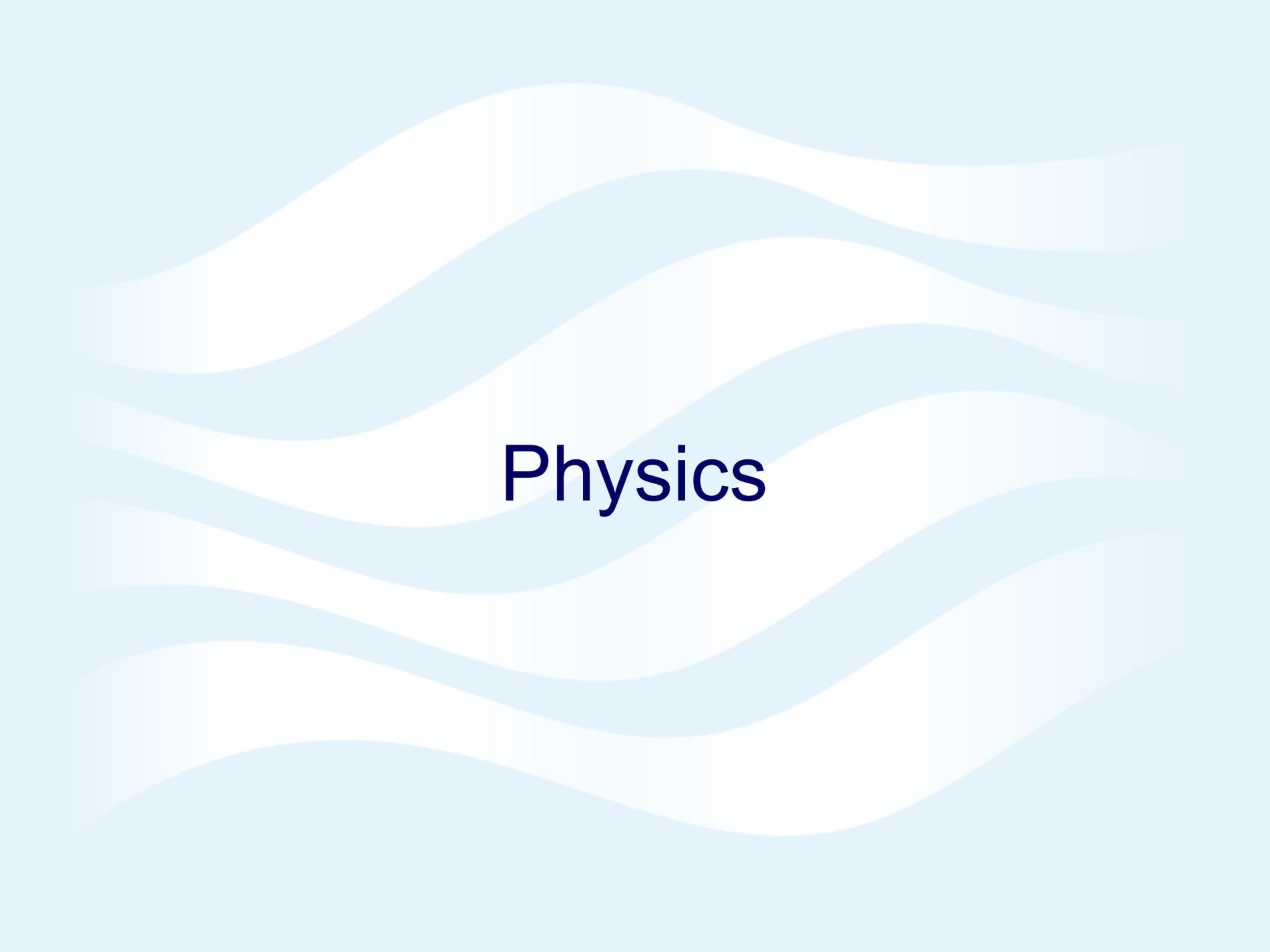
VERTICAL (BL)





0 300 600 900 1200 1500 1800

Visibility
100m orography contours

The background of the slide features a series of horizontal, wavy bands in various shades of light blue and white, creating a fluid, wave-like pattern. The word "Physics" is centered in a dark blue, bold font.

Physics

Physics parametrisations in the Unified Model are designed for a wide range of grid scales, from Mesoscale (12 km) to Climate (~300 km).

Physics formulation needs optimising for high resolution fog forecasting.

➤ Radiation

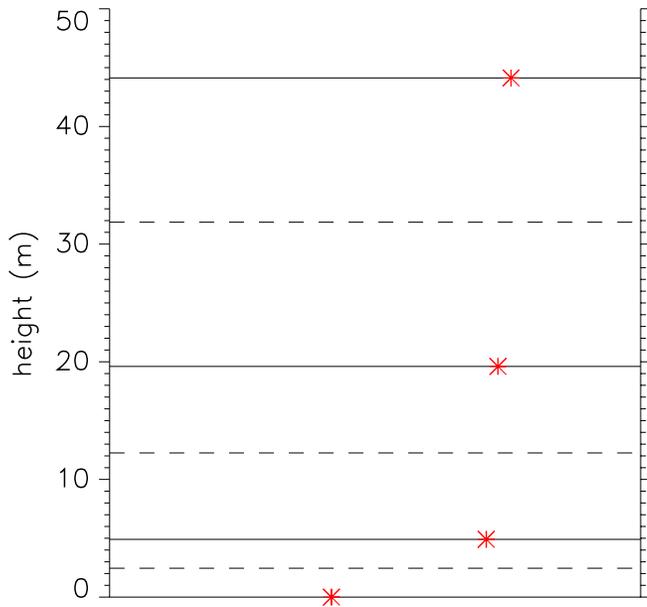
- Temperature profile near surface

➤ Boundary Layer Scheme

- Stability functions

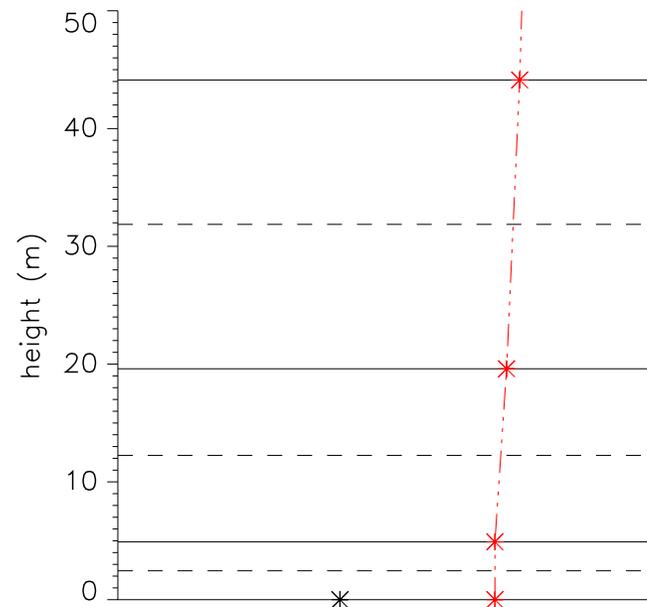
VERTICAL GRID

T on model levels



RADIATION SCHEME

T profile

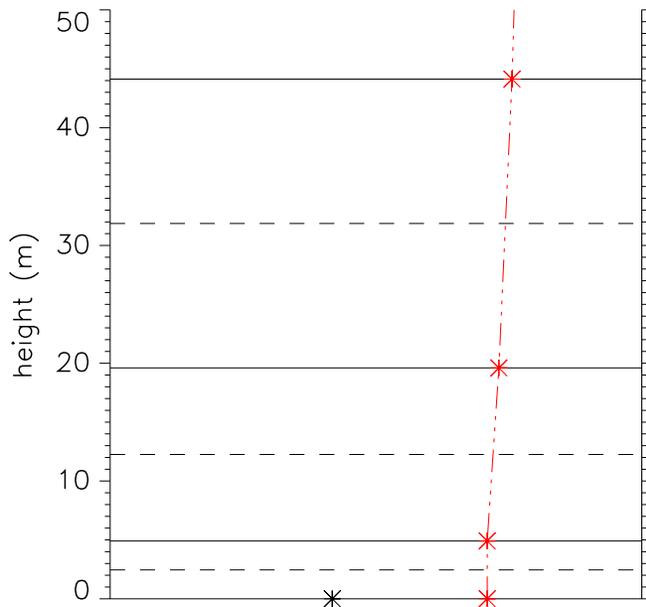


TEMPERATURE PROFILE near surface

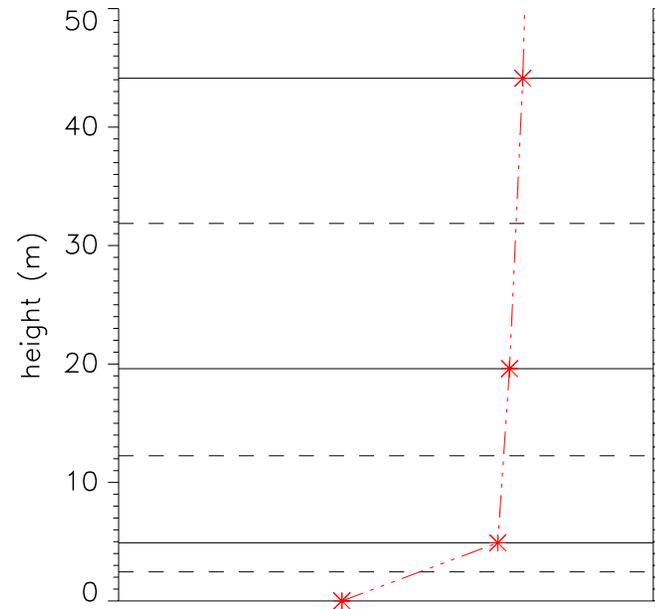
- Radiation scheme sets T at bottom of atmosphere, i.e. at surface, to T on first model theta-level.

RADIATION SCHEME

T profile



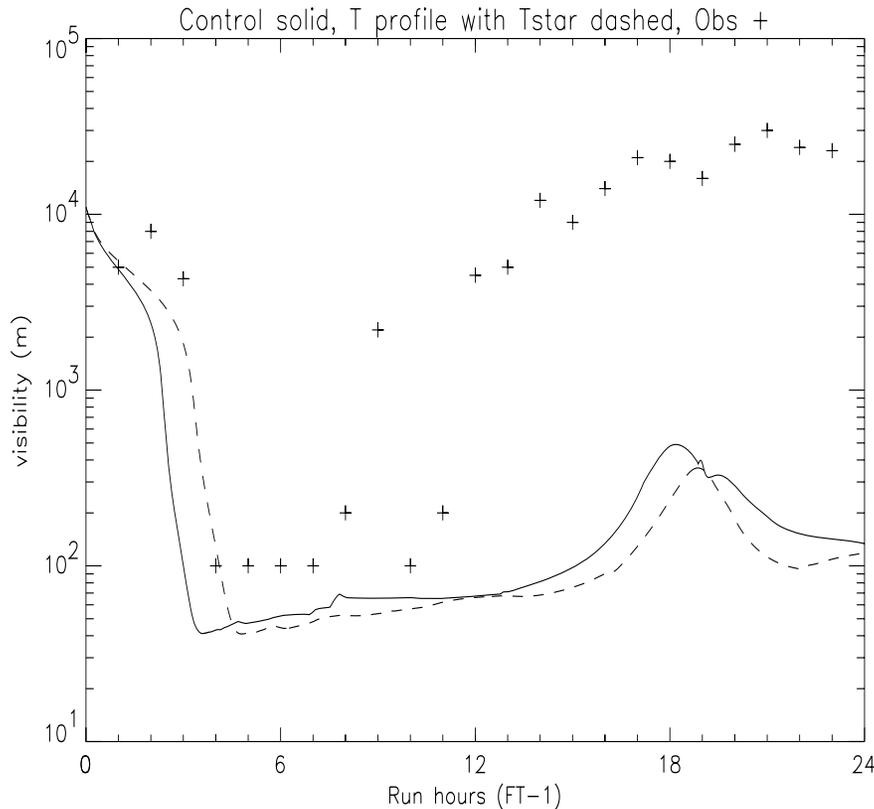
T profile using T_*



TEMPERATURE PROFILE near surface

- Radiation scheme sets T at bottom of atmosphere, i.e. at surface, to T on first model theta-level.
- T_* could be a better choice over land.

Visibility at 1.5m



Using T_* in radiation code profile gives

- Less cooling
- Sharper inversion
- Later fog formation

Why?

Single Column UM 6.0, 76 levels,
Initial profile from 3D model

First order SBL scheme in UM:

$$\overline{\chi w} = -K_{\chi} \frac{\partial \chi}{\partial z}; \quad \chi = U, V, \theta$$

$$K_{\chi} = \lambda^2 \left| \frac{\partial \mathbf{u}}{\partial z} \right| f_{\chi}(Ri)$$

λ Mixing length
 $f_{\chi}(Ri)$ Stability function

$$\frac{1}{\lambda} = \frac{1}{\kappa(z+z_0)} + \frac{1}{\lambda_0}; \quad \lambda_0 = \text{MAX}(40, 0.15h)$$

$$Ri = \frac{\frac{g}{\theta_0} \frac{\partial \theta}{\partial z}}{\left(\frac{\partial U}{\partial z} \right)^2 + \left(\frac{\partial V}{\partial z} \right)^2}$$

Gradient Richardson number
 $Ri < 0.25$ turbulence 'turns on'
(empirically, not in model)

LONG

$$f_{\chi}(Ri) = \frac{1}{1 + 10Ri}$$

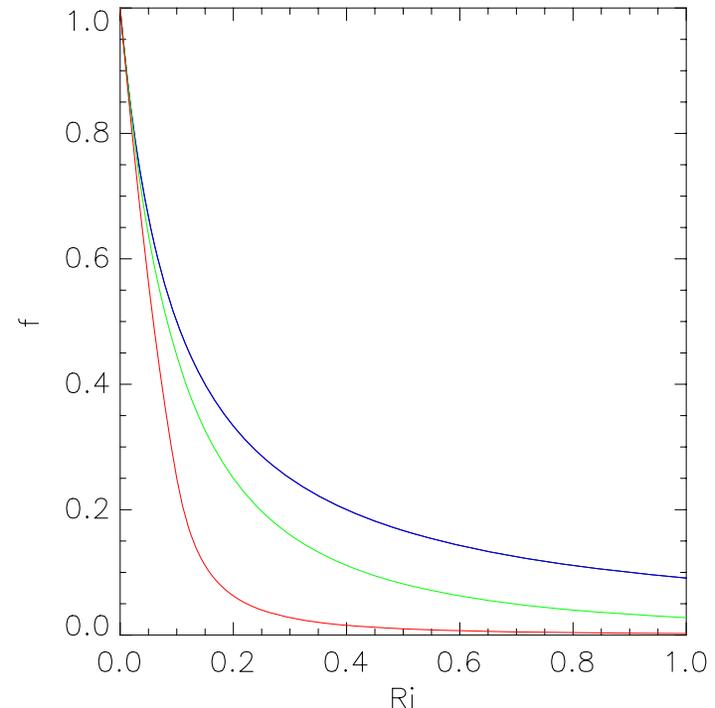
LOUIS

$$f_{\chi}(Ri) = \frac{1}{(1 + 5Ri)^2}$$

SHARPEST

$$f_{\chi}(Ri) = (1 - 5Ri)^2 \quad 0 < Ri < 0.1$$

$$f_{\chi}(Ri) = \frac{1}{(20Ri)^2} \quad Ri > 0.1$$



SHARPEST has least mixing

LONG

$$f_{\chi}(Ri) = \frac{1}{1 + 10Ri}$$

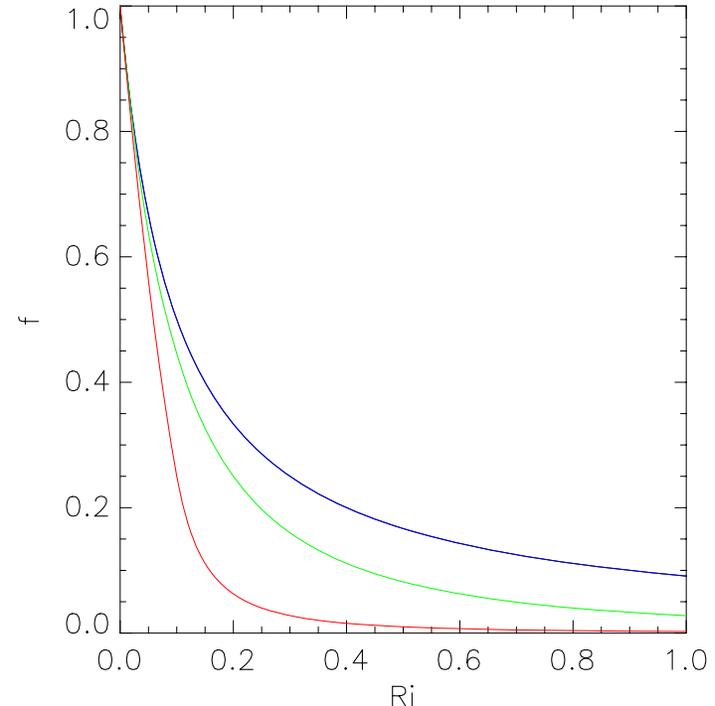
LOUIS

$$f_{\chi}(Ri) = \frac{1}{(1 + 5Ri)^2}$$

SHARPEST

$$f_{\chi}(Ri) = (1 - 5Ri)^2 \quad 0 < Ri < 0.1$$

$$f_{\chi}(Ri) = \frac{1}{(20Ri)^2} \quad Ri > 0.1$$



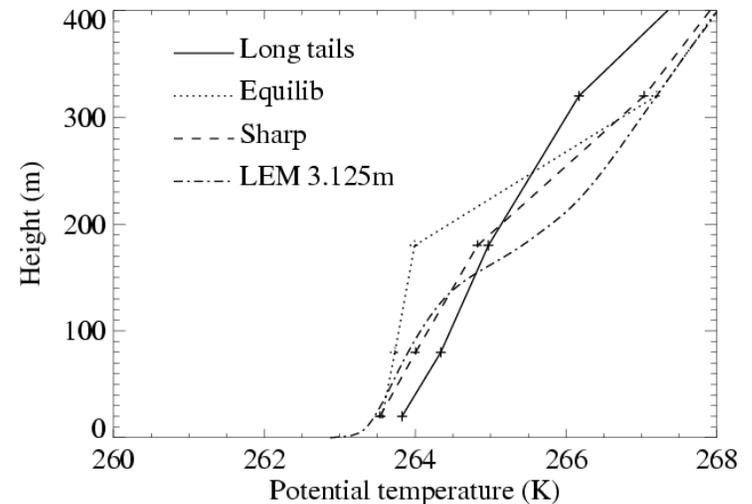
SHARPEST has least mixing

■ UM is tuned!

SHARPEST above 200m, scales to Louis with z below 200m.

- Stability functions with more mixing compensate for sub-grid heterogeneity of surface.
- But SHARPEST is in fashion – best agreement with LES in GABLS. www.gabls.org
- Sub-grid surface more homogeneous at very high resolution.
- Rerun 3D case studies with SHARPEST.

Single Column Model vs LEM



Anne McCabe
Bob Beare

In 3 case studies, model forecasts using SHARPEST stability functions consistently show

- Screen temperature 0.5-1.0 K cooler
- Relative humidity 1-3% higher
- 10m winds 0.2-0.4 ms⁻¹ less
- Lower visibility

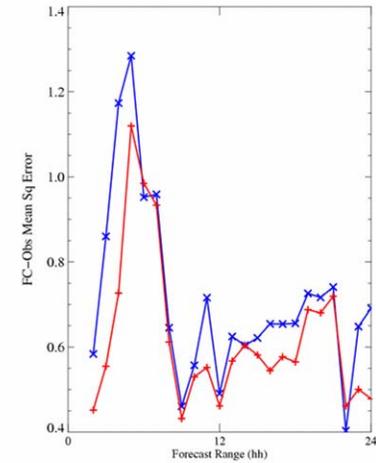
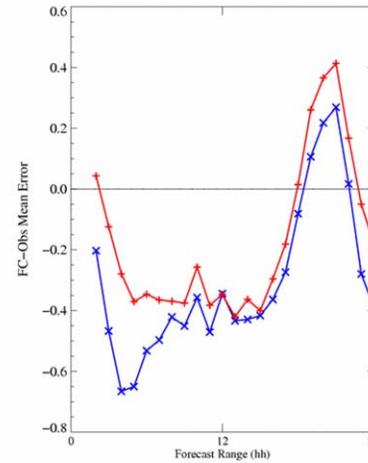
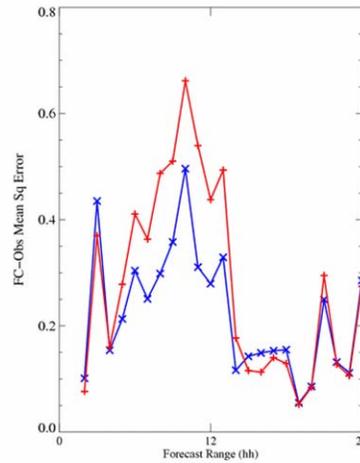
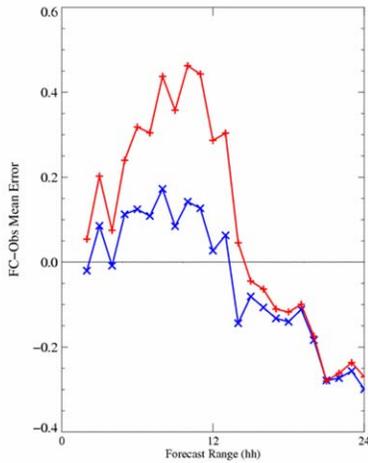
SHARPEST stability functions produce less mixing and hence quicker fog formation, thicker fog and slower fog dissipation.

Impact on forecast varies – best verification is for more patchy fog, e.g. 06-07.11.2003.

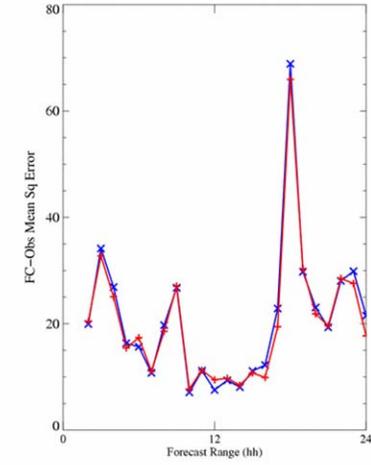
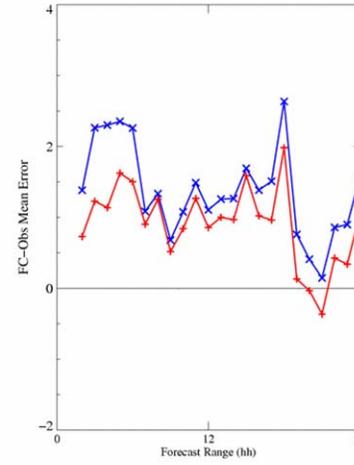
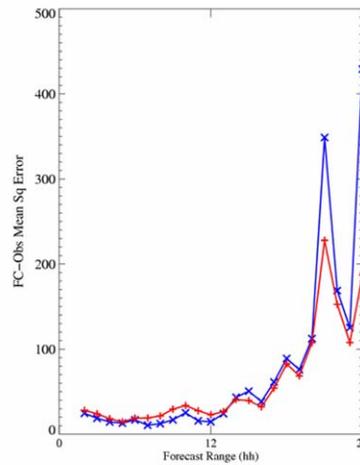
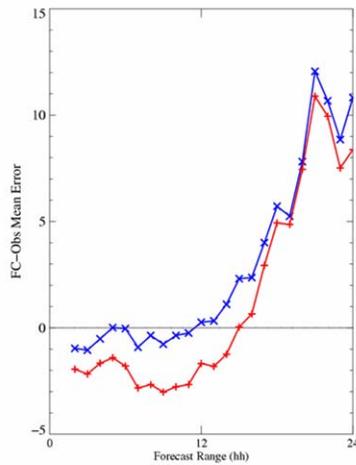
1kL76 (red) vs. SHARPEST (blue) From 18 UTC 06/11/2003

From 18 UTC 09/12/2003

log VIS



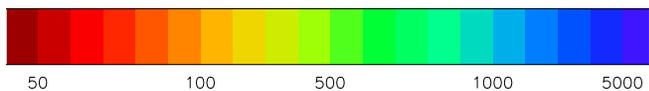
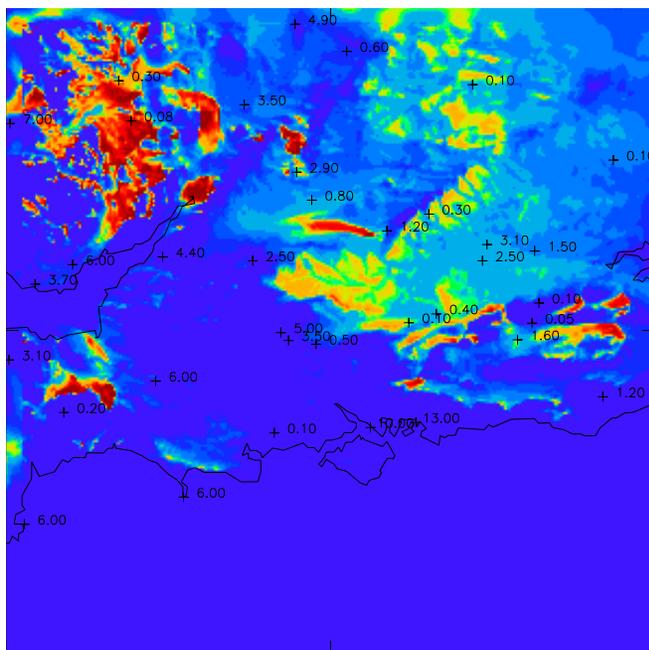
RH



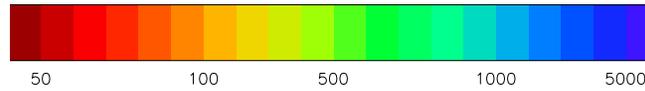
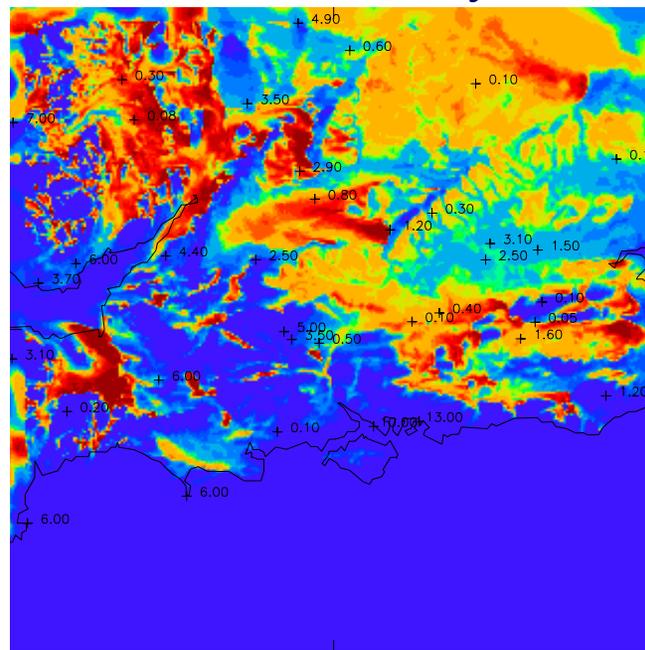
At 04 UTC 07/11/2003 from 18 UTC 06/11/2003 T+10 Forecast



1km L76



SHARPEST stability fns



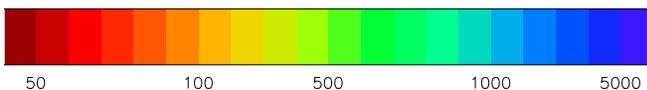
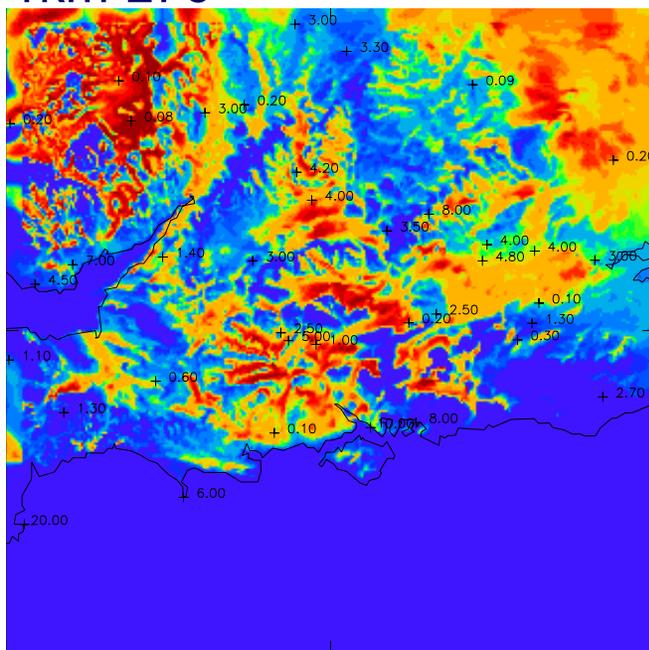
Visibility (m) at station height, synoptic observations (km)

- Radiation scheme – Profile using T_* forms fog later.
- SHARPEST stability functions produce less mixing and hence quicker fog formation, thicker fog and slower fog dissipation. They verify best in patchy fog cases.
- More physics could be investigated, e.g. partial radiation time-stepping, droplet settling.
- Particular attention needed to timing of fog formation and dissipation.
- Use SHARPEST stability functions and Radiation T profile with T_* .

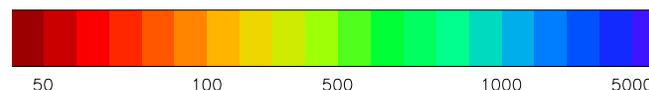
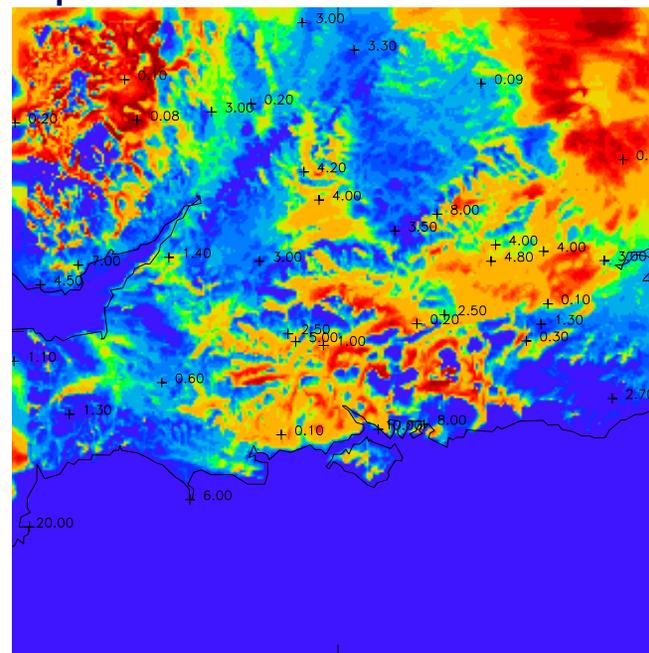
At 21 UTC 09/12/2003 from 18 UTC 09/12/2003 T+3 Forecast



1km L76



SHARPEST stability fns and T profile with T_*



Visibility (m) at station height, synoptic observations (km)

Local Stand-Alone Model

- **Unified Model at Very High Resolution**
 - 1km horizontal grid, 76 vertical levels, small (e.g. 50x50km) domain

- **Model Physics**
 - Optimised for fog forecasting at very high resolution
 - SHARPEST stability functions
 - T profile including T_* in radiation code

- **Single Profile Forcing**
 - Choice of forcing data – model or sonde

- **Run on local PC – using Ported Unified Model**



Initialize 3D model over high resolution domain with

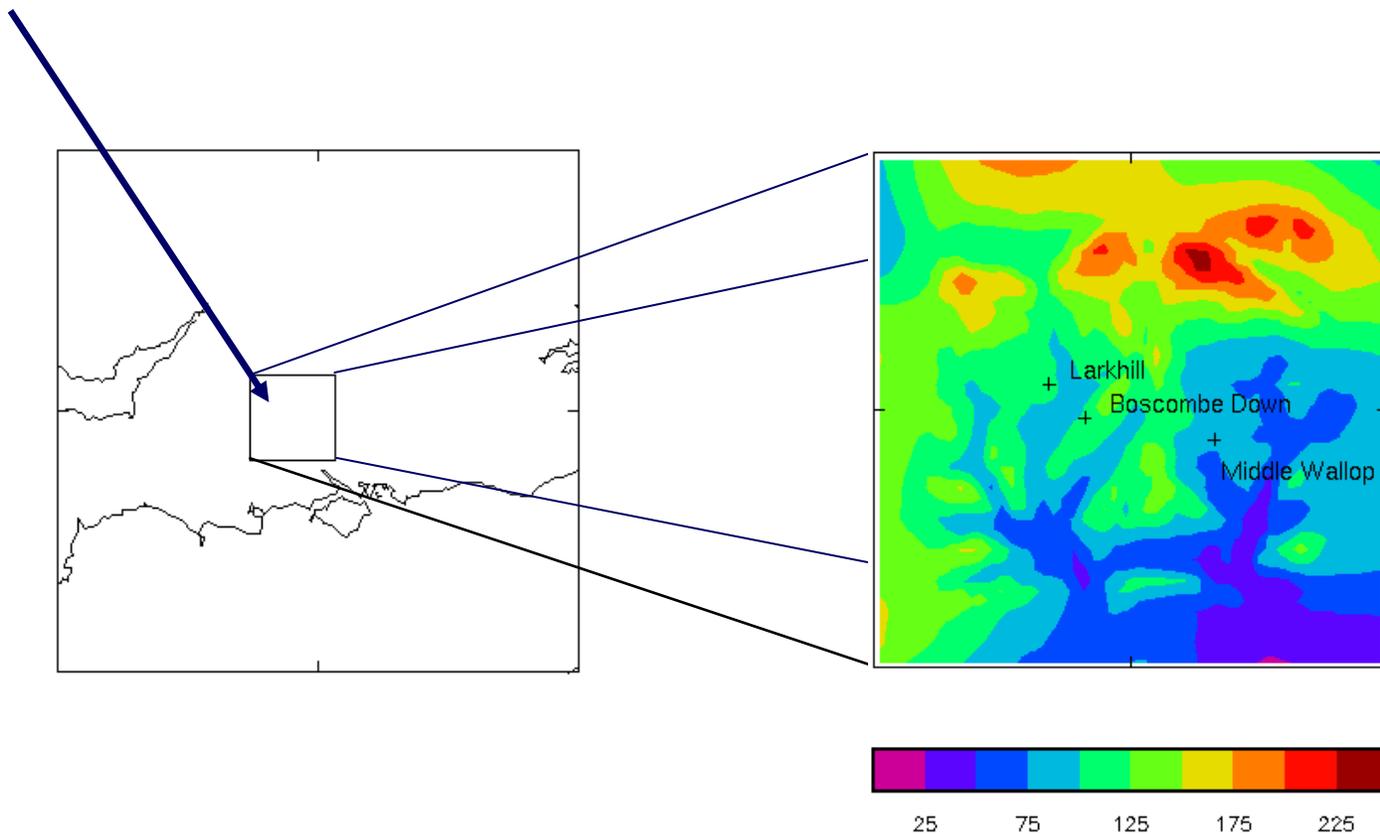
- Profile of θ , q , u and v
- pmsl

Forced hourly with profile through LBCs – option for time-varying profile with model data

Data can be taken from

- [Crisis Area] Mesoscale or Global Model
 - Average
 - Lowest Grid Point
- Local Radiosonde

50 x 50 km domain inset on 300 x 300 km domain

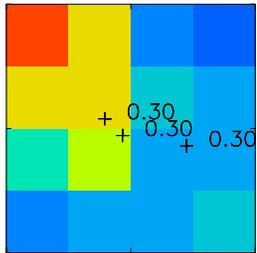


Model orography (m)

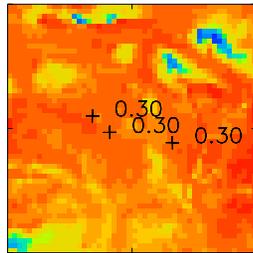
At 07 UTC 03/12/2004 from 18 UTC 02/12/2004 T+13 Forecast - Visibility



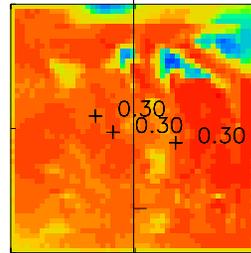
MES



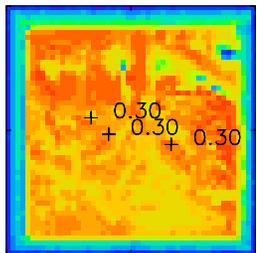
300 x 300 LBC



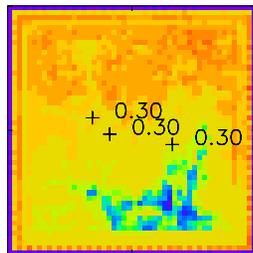
50 x 50 LBC



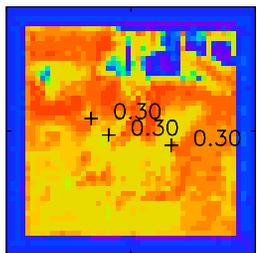
SPF AVE hourly



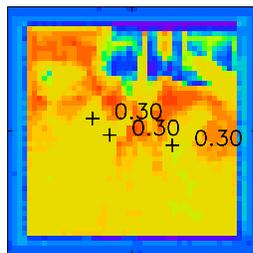
SPF LGP hourly



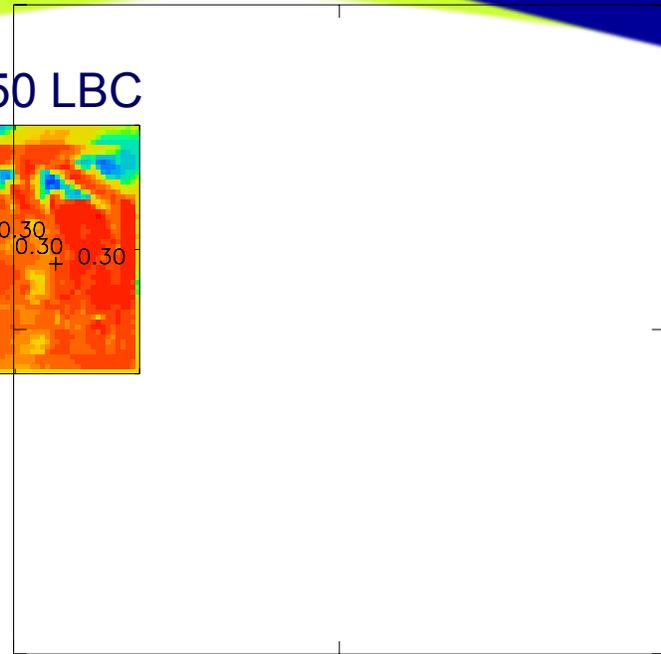
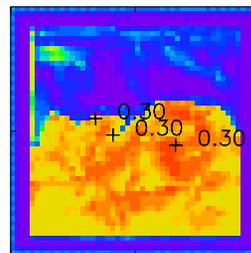
SPF AVE INIT



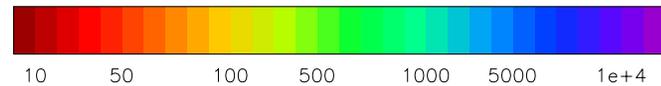
SPF LGP INIT



SPF SONDE



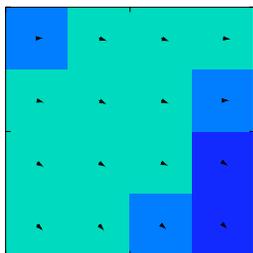
Visibility (m)



At 07 UTC 03/12/2004 from 18 UTC 02/12/2004 T+13 Forecast – 10m Winds

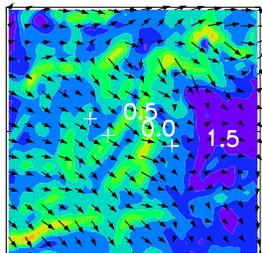


MES



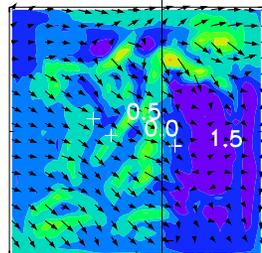
→ 5

300 x 300 LBC



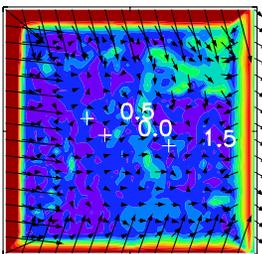
→ 2.5

50 x 50 LBC



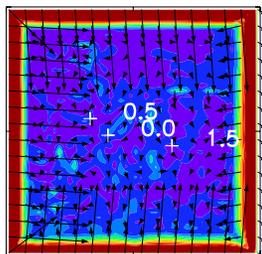
→ 2.5

SPF AVE hourly



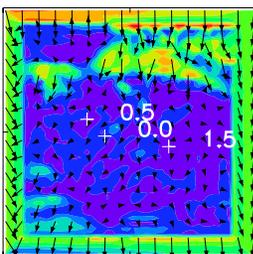
→ 2.5

SPF LGP hourly



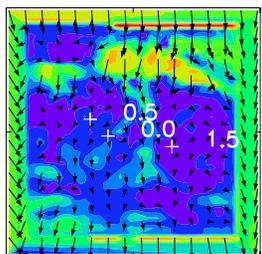
→ 2.5

SPF AVE INIT



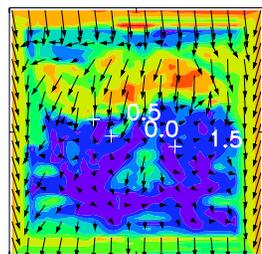
→ 2.5

SPF LGP INIT

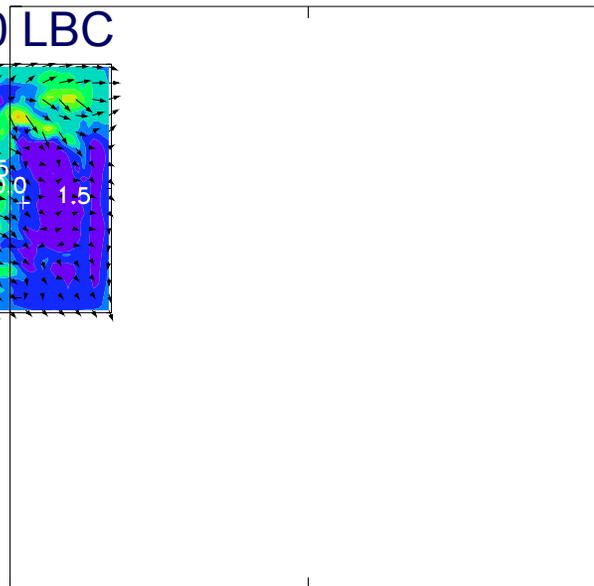


→ 2.5

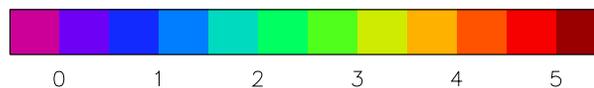
SPF SONDE



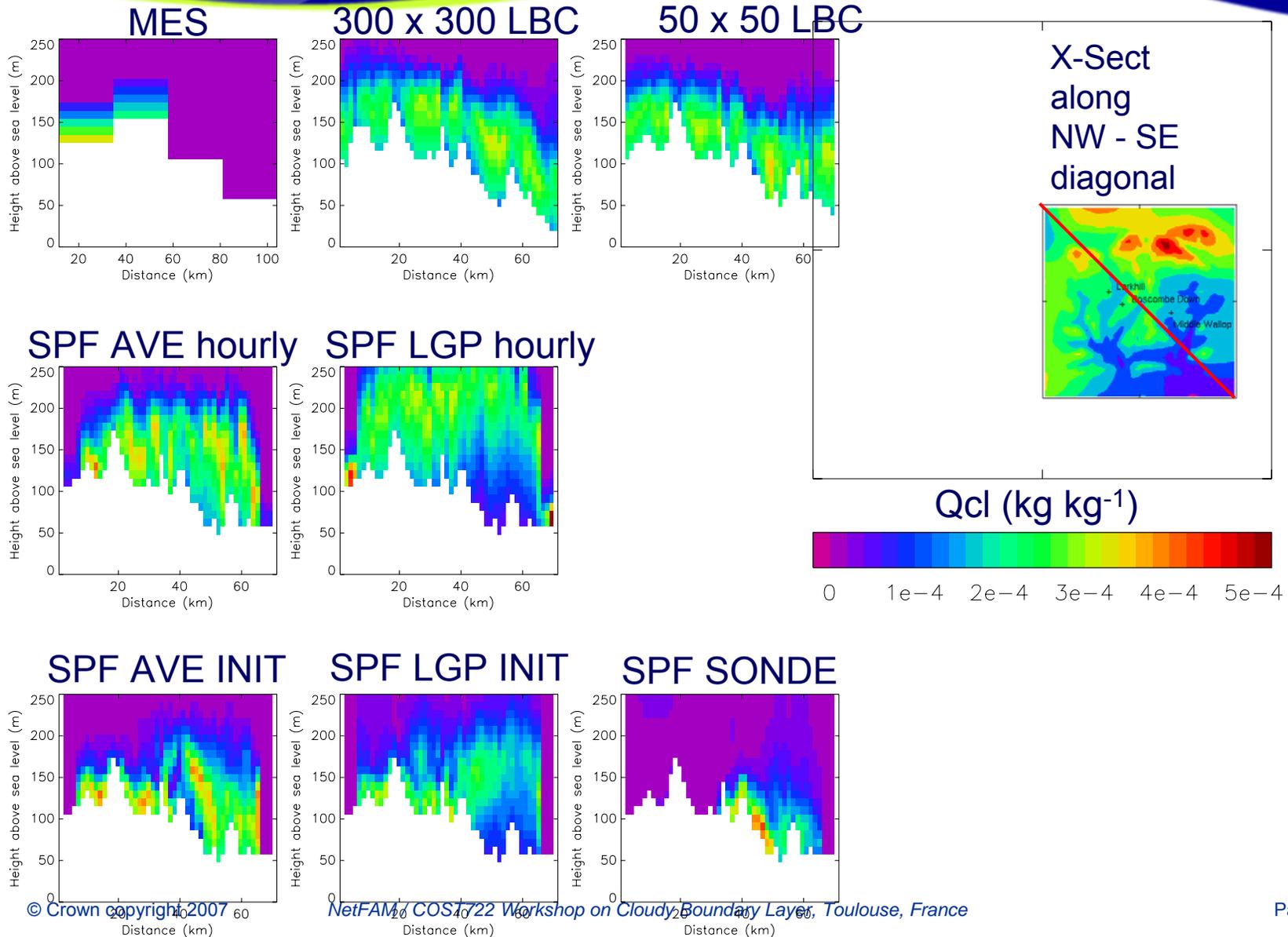
→ 2.5



Wind speed (ms⁻¹)



At 07 UTC 03/12/2004 from 18 UTC 02/12/2004 T+13 Forecast – Cloud Liquid Water X-Sect



Single Profile Forcing at Very High Resolution



- Works remarkably well for radiation fog situations
- Forecast very sensitive to forcing – ensemble approach could be useful
- SONDE as good as any other technique – may allow more recent observational data than model, which will take up to 3 hrs to run and transmit to field

Questions

- *Turton and Brown 1987*, QJRMS **113**, pp. 37-54
- *Roach et al 1976*, QJRMS **102**, pp. 313-333
- *Brown and Roach 1976*, QJRMS **102**, pp. 335-354
- *Capon 2004*, NWP Technical Report FR444

http://www.metoffice.gov.uk/research/nwp/publications/papers/technical_reports/